Sapphire mirror suspension control of KAGRA

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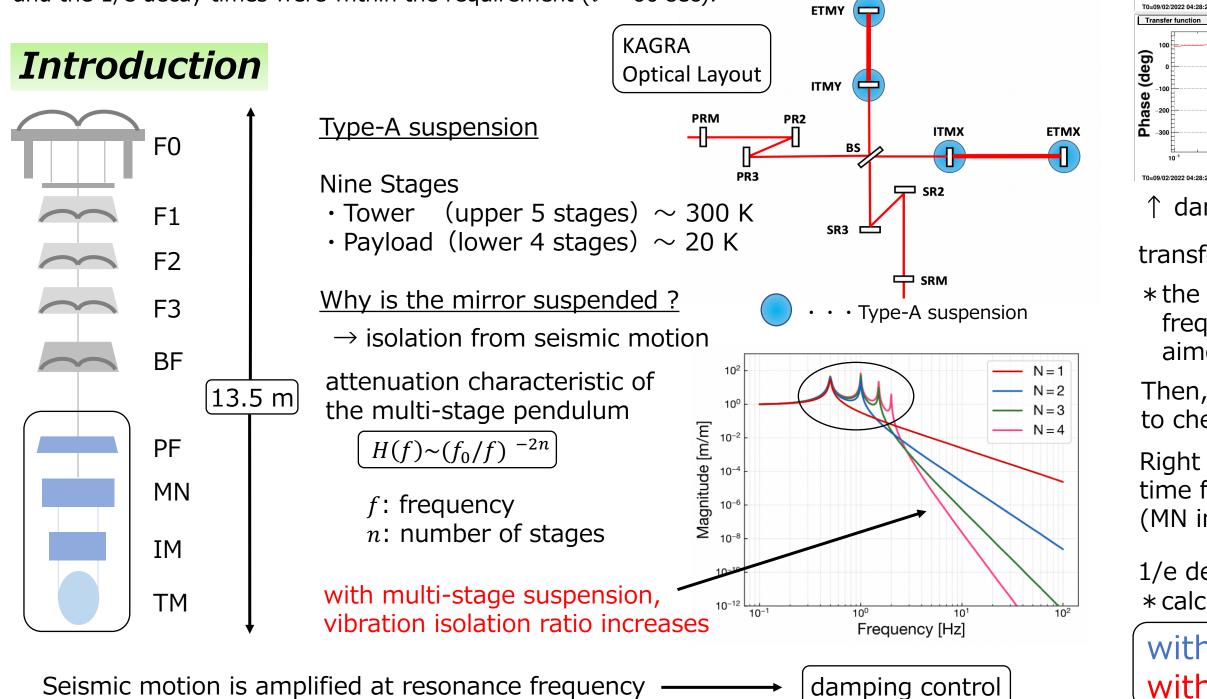
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Abstract

KAGRA, which is based on a Michelson interferometer and has an optical cavity in each arm, needs seismic noise isolation system for mirrors to reach the targeted sensitivity especially for the low frequency range below ~ 100 Hz. To obtain enough seismic noise isolation, a multi-stage suspension system is adopted.

However, the excess motions of the mirror at many resonance frequencies of the multi-stage suspension system are expected without damping controls. These motions should be damped below the requirement levels for the stable interferometer operation as a gravitational wave detector. Then, the requirement for the motion of sapphire mirrors, that are suspended by so called Type-A suspension and form arm Fabry-Perot cavities, is that the 1/e decay time (τ) of the major resonances must be less than 60 sec.

If τ is short, we can reduce dead times during KAGRA operation after earthquakes happen or we accidentally excite motion of any suspension stages. This means that we can quickly recover the interferometer to the observation mode and to lengthen the observation time. To satisfy this requirement for damping motions, we made damping control filters and applied. Then, we checked whether the resonance peaks are suppressed or not, and the 1/e decay times were within the requirement ($\tau = 60$ sec).

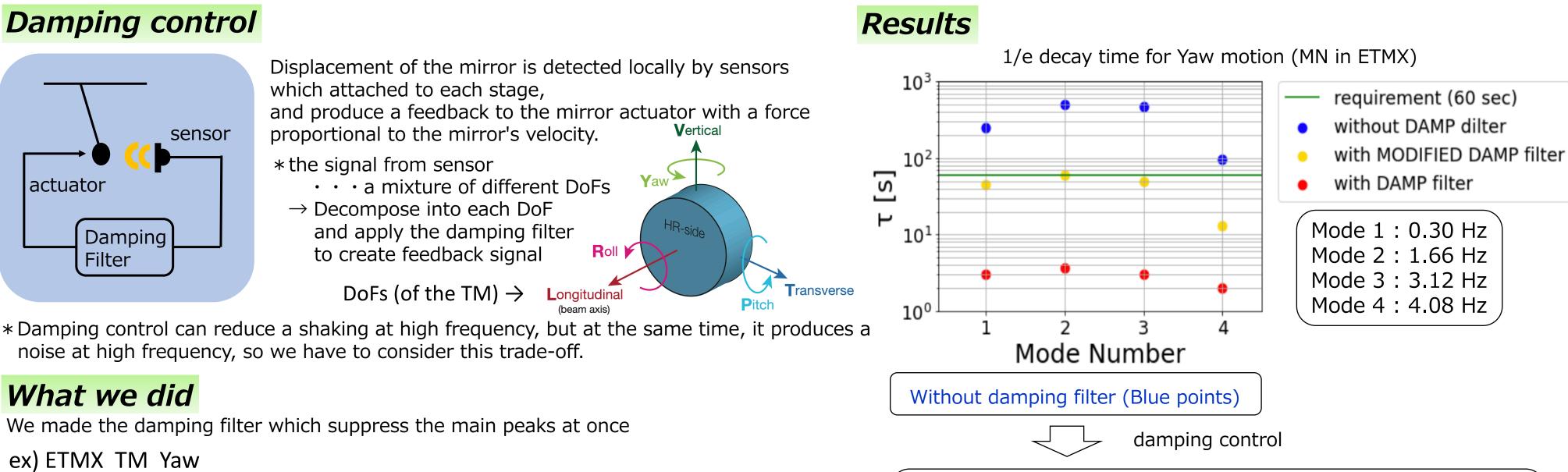




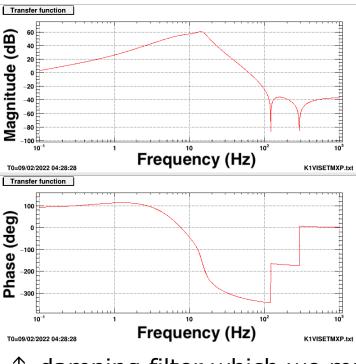
* Damping control can reduce a shaking at high frequency, but at the same time, it produces a noise at high frequency, so we have to consider this trade-off.



We made the damping filter which suppress the main peaks at once



What we did



damping filter which we made

transfer function (ETMX MN Y) \rightarrow

* the peaks at the resonant frequencies were damped as we aimed.

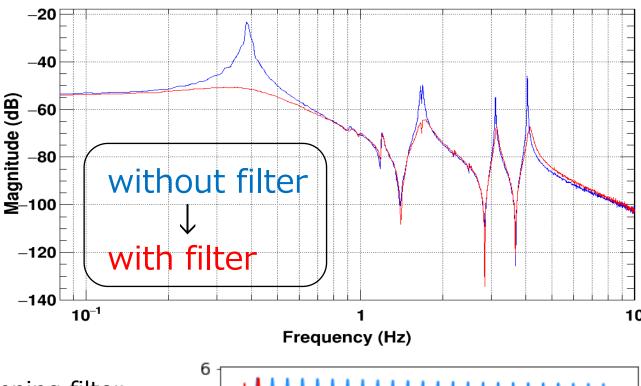
Then, we measured 1/e decay time to check the performance of the damping filter

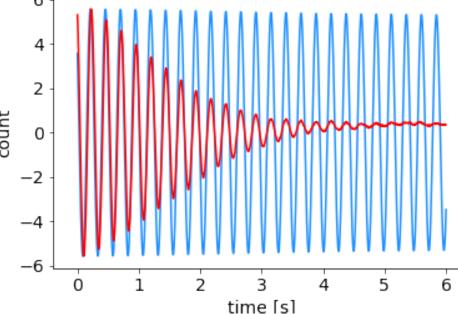
Right figure shows the difference of the decay time for 4th mode (4.08 Hz) of Yaw motion (MN in ETMX)

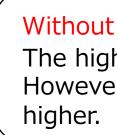
1/e decay time (requirement = 60 s) * calculated by fitting with $y = a \exp(-t/b) + c$

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without damping filter \cdot \cdot \cdot 97.1 \pm 0.4 s
with damping filter \cdot \cdot \cdot 2.07 \pm 0.21 s
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- 1. Apply a filter that is proportional to frequency
- 2. Apply a low-pass filter from a certain frequency (to reduce the noise at high frequencies)
- 3. Determine the gain so that the UGF has a phase margin







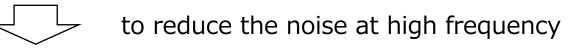


Summary and Future steps

- control noise itself.

Without damping filter (Red points)

The higher the gain of the damping filter, the shorter τ becomes. However, at the same time, the noise at high frequencies becomes



Final results (Yellow points)

We could design a filter that satisfies the requirement for 1/e decay time and keeps the noise as low as possible.

• We made the damping filter for Yaw action of MN in ETMX and the requirement for the 1/e decay time was satisfied. We also could adjust the gain of the damping filter to minimize the excitation due to the

• This allowed us to quickly reduce the shaking when an earthquake occurred or when the suspension was accidentally shaken. This mean that the interferometer could be quickly returned to its locked state, which would lead to longer observation times.

• For the next step, we will do the same process for the other degrees of freedom (Longitudinal, Transverse, Roll, Pitch) and also for the other suspensions (ETMX, ITTMX, ITMY).

• Consider whether modal dumping is necessary.