Suspension Control in KAGRA - Gravitational Wave Detector -Masahide TAMAKI (e-mail: tamaki83@icrr.u-tokyo.ac.jp) D2 (Miyoki Laboratory) / Gravitational-Wave group

ICRR Master & Doctor Thesis Workshop (21. Feb)



Gravitational Wave

Ripples of



BS

laser room

Suspension for

Particularly important characteristics of KAGRA

These are two unique features not found in other GW detectors



space-time transmitted at the speed of light

 Mainly observe GW from compact binary coalescences

black hole, neutron star....



Laser interferometer

3 km : use the laser to measure changes in mirror spacing due to GW

main mirror (13.5 m & 9 stages)

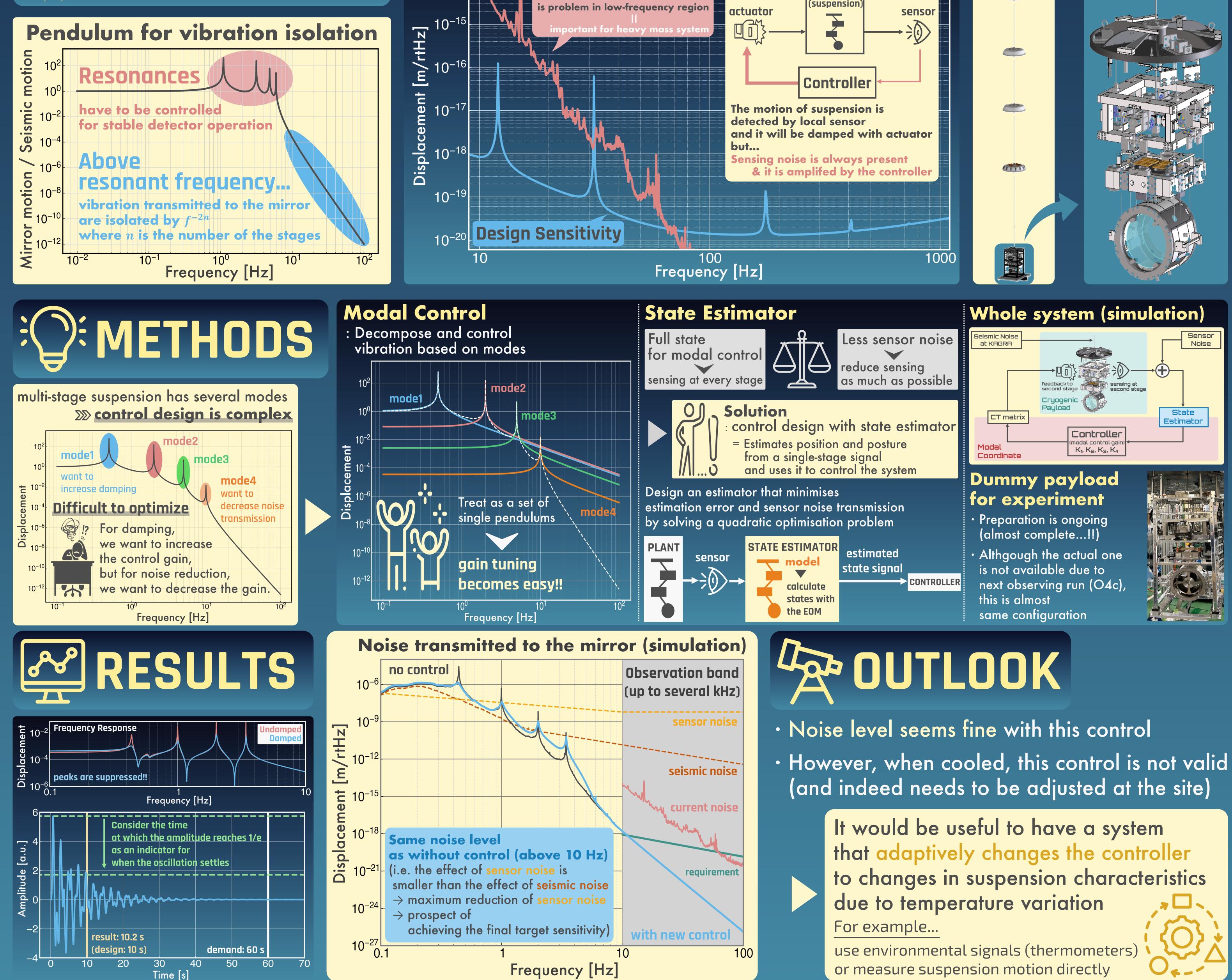
To reduce seismic noise, detector is constructed underground

where ground vibrations are small (Kamioka-cho, Hida-city, Gifu-prefecture)

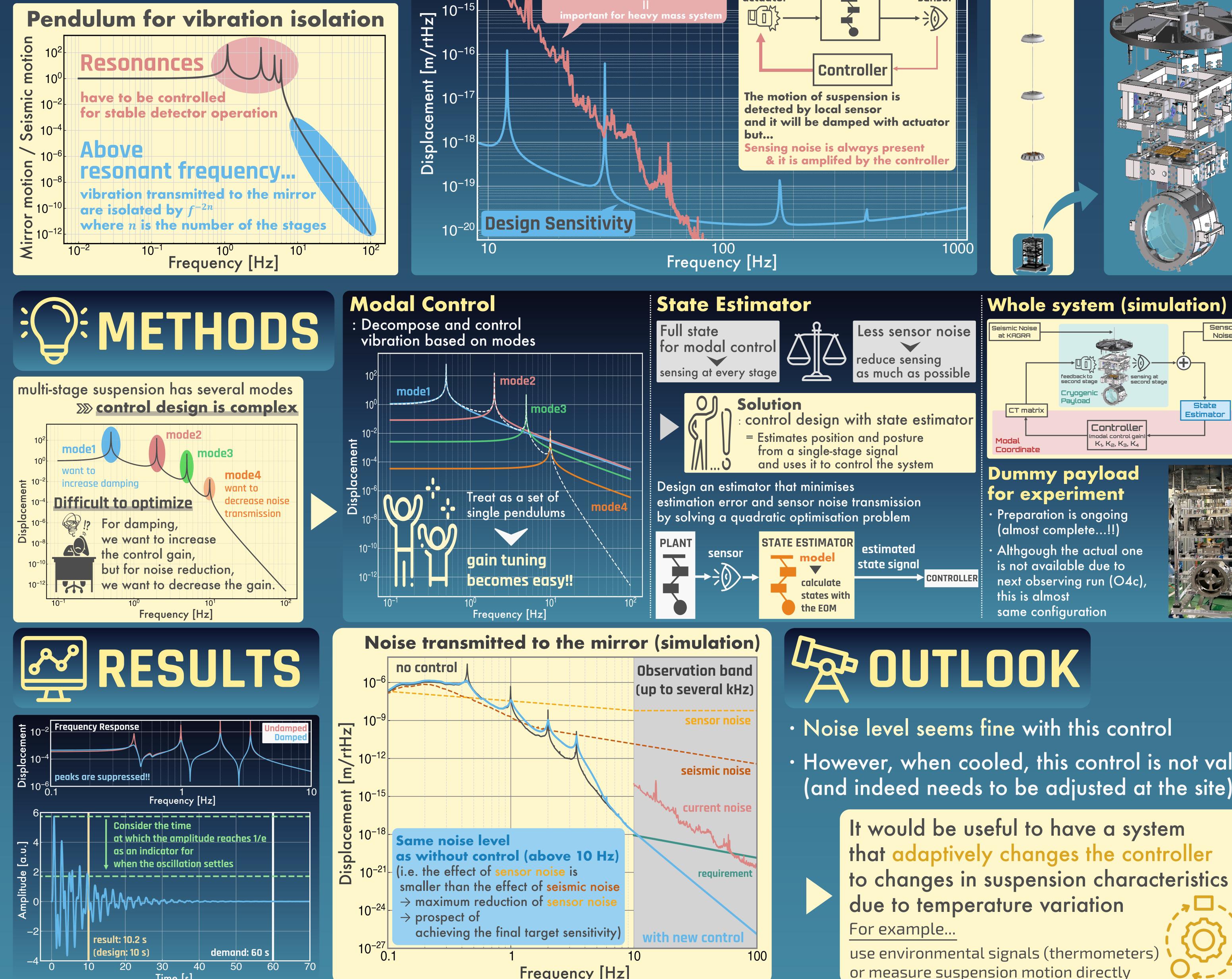
Cryogenic

To reduce thermal noise, main sapphire mirror is cooled down to about 20 K (sapphire has less thermal conduction and mechanical loss when cooled to cryogenic temperature)

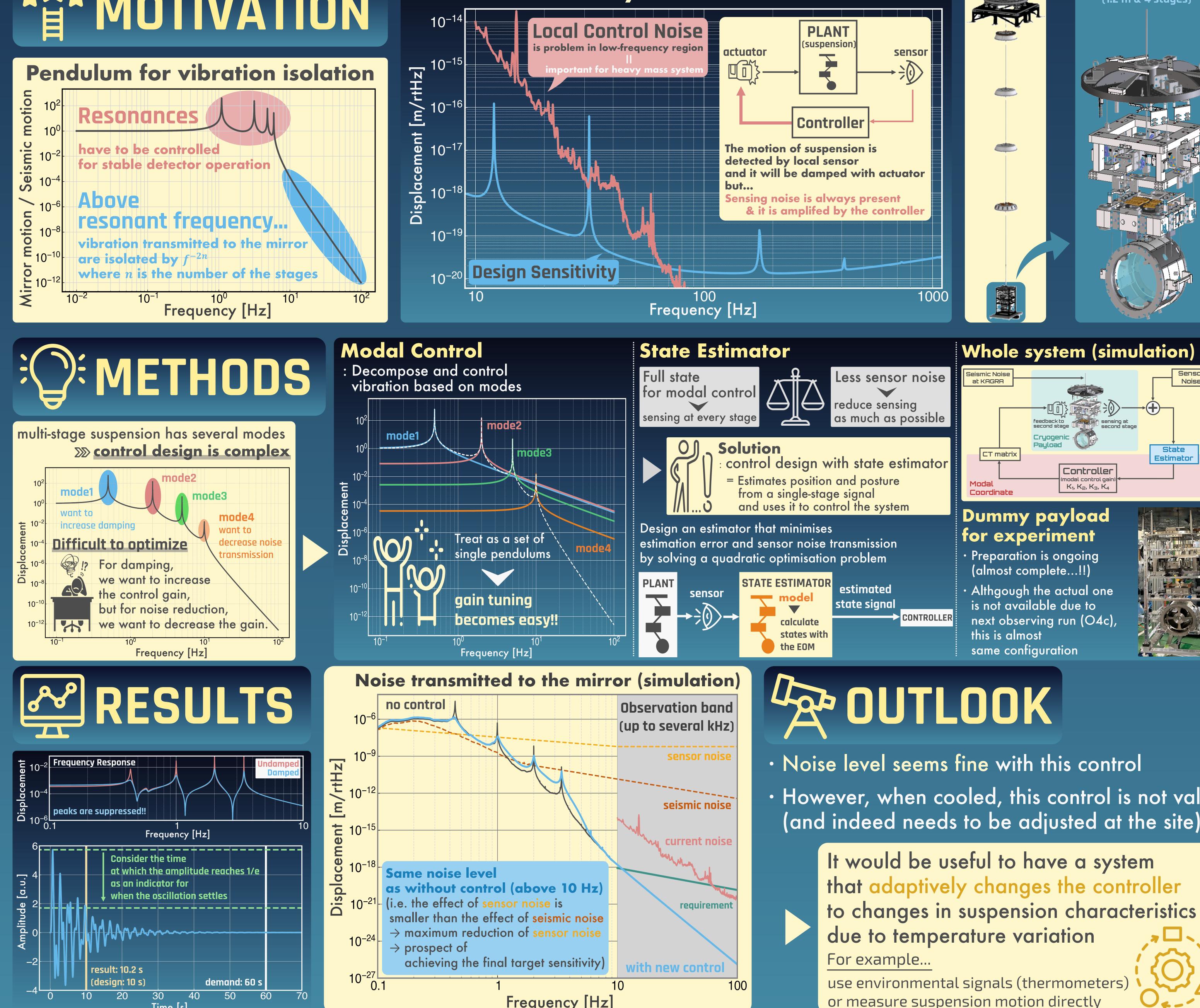
> Cryogenic payload (1.2 m & 4 stages)







KAGRA Sensitivity and Local Control Noise



or measure suspension motion directly